

**Barium isotopic compositions of ordinary chondrites.** K. Misawa<sup>1,2</sup>, Tatsunori Yokoyama<sup>3</sup>, and S. Yoneda<sup>3</sup>, <sup>1</sup>National Institute of Polar Research, 10-3 Midoricho, Tachikawa, 190-8518, Japan (misawa@nipr.ac.jp), <sup>2</sup>SOKENDAI, <sup>3</sup>Department of Science and Engineering, National Museum of Nature and Science, 4-1-1 Amakubo, Tsukuba, 305-0005, Japan.

**Introduction:** Relative to the Sun's photosphere, moderately volatile elements are depleted in the Earth, Moon, Mars, Vesta and all meteorites except CI-chondrites. In relatively few cases, very alkali-rich materials have been observed in chondritic breccias. Previous studies revealed that alkali elements in Krähenberg (LL5), Bhola (LL3–6), Y-74442 (LL4), and Acfer 111 (H3–6) fragments are enriched and fractionated relative to CI-chondrites (Fig. 1) with heavier alkalis being more enriched ( $C_{S_{CI-norm}} > Rb_{CI-norm} > K_{CI-norm}$ ) [1–4]. Cesium-135 ( $t_{1/2} = 2.3$  Myr) is a short-lived nuclide that can date early solar system events [5–9]. If  $^{135}Cs$  was present in the early solar system, we can detect a  $^{135}Ba$  excess in a reservoir having a high Cs/Ba ratio. In this study, we focus on the  $^{135}Cs$ - $^{135}Ba$  system of rock fragments in chondritic breccias to better understand the extent and timing of the heavy alkali enrichments in the early solar system.

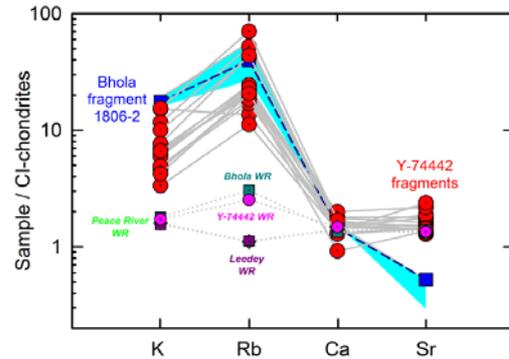
**Experimental:** The Ba isotopic data were obtained on a TIMS at NMNS by a static multidynamic mode utilizing the zoom lens capability. Instrumental mass fractionation was corrected using the exponential law with  $^{134}Ba/^{136}Ba = 0.3078$  as the normalizing ratio. A single Ba isotopic analysis usually consisted of 540 cycles that were averaged. Possible isobaric interferences of  $^{138}La$  and  $^{136,138}Ce$  were monitored and corrected using  $^{139}La$  and  $^{140}Ce$  assumed natural  $^{138}La/^{139}La$  and  $^{136,138}Ce/^{140}Ce$  ratios, which was always negligible. Two Ba standards (SPEX ICP-MS standard and JM Alfa Aesar, Suprapur) as well as whole-rock samples of the Leedeey (L6) chondrite were analyzed. All data are presented as  $\mu^{13x}Ba$  values, which are the parts per million deviations from the standard:  

$$\mu^{13x}Ba = [(^{13x}Ba/^{136}Ba)_{sam}/(^{13x}Ba/^{136}Ba)_{std} - 1] \times 10^6.$$

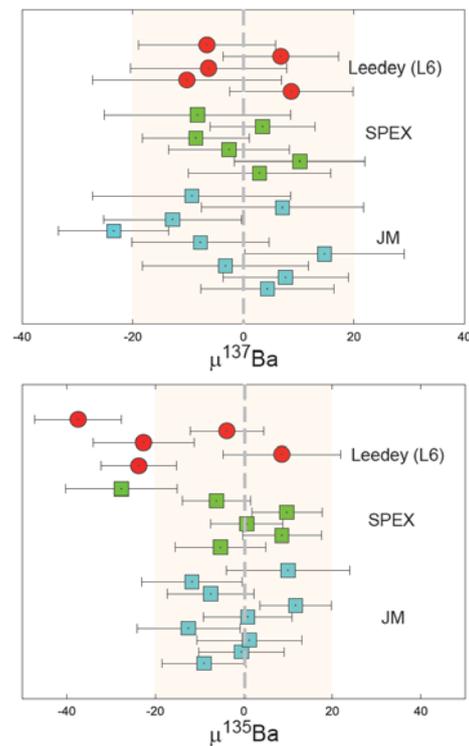
**Results and Discussion:** The Ba isotopic data are shown in Fig. 2. External precisions of  $^{135}Ba/^{136}Ba$  and  $^{137}Ba/^{136}Ba$  ratios of the standards (50 ng of Ba) are  $\sim 20$  ppm ( $2\sigma$ ) (solid squares). The  $^{135}Ba/^{136}Ba$  and  $^{137}Ba/^{136}Ba$  ratios of whole-rock samples of Leedeey (L6) are normal within the errors (Fig. 2, solid circles). The result is consistent with the previous studies: the nucleosynthetic isotopic effects,  $r$ -process contributions to the  $^{135,137}Ba$  excesses, are smaller in ordinary chondrites than in several CM chondrites [5–9].

The Ba isotopic composition of the spiked sample (composite  $^{40}K$ - $^{48}Ca$  and  $^{87}Rb$ - $^{84}Sr$  spikes) of Leedeey (L6) was clearly different from those of standard, indicating a contribution of Ba in the spikes becomes too large to ignore. The Y-74442 and Bhola samples used for the K-Ca and Rb-Sr isotopic studies [3,4] also showed scattered Ba isotopic

signatures as expected.



**Fig. 1.** CI-normalized alkali and alkaline earth abundances of lithic fragments in the LL-chondritic breccias, Y-74442 and Bhola [4]. Shaded area represents ranges of Krähenberg and Bhola fragments [1].



**Fig. 2.**  $^{137}Ba/^{136}Ba$  (upper) and  $^{135}Ba/^{136}Ba$  (lower) results, normalized to  $^{134}Ba/^{136}Ba = 0.3078$  for standards (squares) and whole-rock samples of Leedeey (circles). Error bars are  $2\sigma_m$ .

**References:** [1] Wlotzka F. *et al.* (1983) *GCA* **47**, 743–757. [2] Wlotzka F. *et al.* (1992) *Meteorit.* **27**, 308. [3] Yokoyama Tatsunori *et al.* (2013) *EPSL* **366**, 38–48. [4] Yokoyama Tatsunori *et al.* (2015) *LPSC* **46**, #1695. [5] Hidaka H. *et al.* (2001) *EPSL* **103**, 459–466. [6] Hidaka H. *et al.* (2003) *EPSL* **214**, 455–466. [7] Qin L. *et al.* (2011) *GCA* **75**, 7806–7828. [8] Hidaka H. & Yoneda S. (2013) *Sci. Rep.* **3**, 1330. [9] Bermingham K.R. *et al.* (2014) *GCA* **133**, 463–478.